15th European Workshop on String Theory Zürich, September 9th, 2009

# Schrödinger black holes with extremal limits



based on arXiv:0907.1892 (with Aninda Sinha) - see also arXiv:0808.1271



AdS/CFT and *experimentally testable systems* 

Applications to *condensed matter systems* 

Non-relativistic superfluid phases and (non)Fermi liquids?

 
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Can string theory be used to access experimentally testable systems?

 
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Can string theory be used to access experimentally testable systems?

The *gauge/string duality* is an invaluable tool to study *strongly coupled* physical systems

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The *gauge/string duality* is an invaluable tool to study *strongly coupled* physical systems



Applications to *strong interactions* (*confinement,*  $\chi$ *SB, quark-gluon plasma,* ...)

Can string theory be used to access experimentally testable systems?

The *gauge/string duality* is an invaluable tool to study *strongly coupled* physical systems



Applications to *strong interactions* (*confinement,*  $\chi$ *SB, quark-gluon plasma,* ...)



Applications to *condensed matter systems* 



The *gauge/string duality* is an invaluable tool to study *strongly coupled* physical systems



Applications to *strong interactions* (*confinement, \chiSB, quark-gluon plasma, ...*)

Applications to *condensed matter systems* 

Many examples of *strongly coupled scale invariant* systems in condensed matter

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Study non-relativistic systems at finite temperature / density

 
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Study non-relativistic systems at finite temperature / density AdS/CFT Black holes with non-relativistic asymptotic symmetry

 
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Study non-relativistic systems at finite temperature / density AdS/CFT Black holes with non-relativistic asymptotic symmetry

IN THIS TALK

Derive and study charged black hole solution dual to non-relativistic scale invariant (2+1)-dimensional conformal field theory admitting an extremal limit (T = 0)

Study non-relativistic systems at finite temperature / density AdS/CFT Black holes with non-relativistic asymptotic symmetry



 
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> Derive and study charged black hole solution dual to non-relativistic scale invariant (2+1)-dimensional conformal field theory admitting an extremal limit (T = 0)

Adding *U(1) charge* may allow spontaneous breaking by a *condensate*  $\Longrightarrow$  *superfluid phase* 

[Gubser, Hartnoll-Herzog-Horowitz, ...]

Near-horizon geometry of charged extremal black holes exhibits an  $AdS_2$  factor that has been shown to be related to the emergence of sharp Fermi surfaces in the dual theory (non-)Fermi liquids

[Liu-McGreevy-Vegh, Faulkner-Liu-McGreevy-Vegh]

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 $k_F$ 

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D PHASE LE TRANSITION  $k_F$ 

Symmetry of non-relativistic "conformal" systems is Schrödinger algebra

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Symmetry of non-relativistic "conformal" systems is Schrödinger algebra

**GALILEAN SYMMETRIES** ROTATIONS  $M_{ij}$  TRANSLATIONS  $H P_i$  BOOSTS  $K_i$ 

 $\operatorname{Sch}_d$ 

 
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**DILATATIONS**  $x_i \rightarrow \lambda x_i \quad t \rightarrow \lambda^2 t$ 



 
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Symmetry of non-relativistic "conformal" systems is Schrödinger algebra

$\operatorname{Sch}_d$	GALILEAN SYMMETRIES	ROTATIONS $M_{ij}$	TRANSLATION	S H P <sub>i</sub>	<b>BOOSTS</b> $K_i$
	DILATATIONS $x_i \rightarrow \lambda x_i$	$t  ightarrow \lambda^2 t$	TIME AND SPAC	E SCALE <b>DII</b>	FFERENTLY
	"SPECIAL CONFORMAL"	$x_i \to \frac{x_i}{1+\mu t}  t \to$	$\frac{t}{1+\mu t}$	+ CENTR	AL ELEMENT





















Obtain *non-relativistic ground state* via *TsT* of  $AdS_5 \times S^5$ 

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AdS<sub>5</sub> × S<sup>5</sup>  
$$ds^{2} = \frac{r^{2}}{l^{2}} \left( -4dx^{+}dx^{-} + d\bar{y}^{2} \right) + l^{2} \left( \left( d\psi + P \right)^{2} + ds_{\mathbb{CP}^{2}}^{2} \right)$$
$$F_{5} = \dots$$

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Obtain *non-relativistic ground state* via TsT of  $AdS_5 \times S^5$ 

AdS<sub>5</sub> × S<sup>5</sup>  
$$ds^{2} = \frac{r^{2}}{l^{2}} \left( -4dx^{+}dx^{-} + d\bar{y}^{2} \right) + l^{2} \left( d\psi + P \right)^{2} + ds^{2}_{\mathbb{CP}^{2}} \right)$$
$$F_{5} = \dots$$
$$S^{5} \text{ AS } U(1) \text{ FIBRATION ON } \mathbb{CP}^{2}$$

 
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Obtain *non-relativistic ground state* via *TsT* of  $AdS_5 \times S^5$ 



 

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Obtain *non-relativistic ground state* via TsT of  $AdS_5 \times S^5$ 

AdS<sub>5</sub> × S<sup>5</sup>  

$$ds^{2} = \frac{r^{2}}{l^{2}} \left(-4dx^{+}dx^{-} + d\vec{y}^{2}\right) + l^{2} \left((d\psi + P)^{2} + ds^{2}_{\mathbb{CP}^{2}}\right)$$

$$F_{5} = \dots$$

$$ds^{2} = \frac{r^{2}}{l^{2}} \left(-4dx^{+}dx^{-} - 4\gamma^{2}r^{2} \left(dx^{+}\right)^{2} + d\vec{y}^{2}\right) + l^{2} \left((d\psi + P)^{2} + ds^{2}_{\mathbb{CP}^{2}}\right)$$

$$F_{5} = \dots$$

$$B = -2\gamma r^{2}dx^{+} \wedge (d\psi + P)$$

 
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$$\begin{array}{c} \operatorname{AdS}_{5} \times S^{5} \\ \operatorname{AdS}_{5} \times S^{5} \\ F_{5} = \ldots \\ \end{array} \\ \begin{array}{c} schr{\"o}dinger \\ space-time \\ F_{5} = \ldots \\ F_{5} = \ldots \\ F_{5} = \ldots \\ \end{array} \\ \begin{array}{c} schr{\`o}dinger \\ space-time \\ F_{5} = \ldots \\ \end{array} \\ \begin{array}{c} schr{\`o}dinger \\ Schr{\`o}dinger \\ space-time \\ \end{array} \\ \begin{array}{c} schr{\`o}dinger \\ Schr{\`o}dinger \\ space-time \\ \end{array} \\ \begin{array}{c} schr{\`o}dinger \\ Schr{\`o}dinger \\ space-time \\ \end{array} \\ \begin{array}{c} schr{\`o}dinger \\ Schr{\`o}dinger \\ space-time \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ Schr{\ro}dinger \\ space-time \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ Schr{\ro}dinger \\ space-time \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ Schr{\ro}dinger \\ space-time \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ Schr{\ro}dinger \\ space-time \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ Schr{\ro}dinger \\ Schr{\ro}dinger \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ Schr{\ro}dinger \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ Schr{\ro}dinger \\ Schr{\ro}dinger \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ Schr{\ro}dinger \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ \end{array} \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ \end{array} \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ \end{array} \\ \end{array} \\ \begin{array}{c} schr{\ro}dinger \\ \end{array} \\ \begin{array}{c$$

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$$\begin{aligned} \operatorname{AdS}_{5} \times S^{5} & ds^{2} = \frac{r^{2}}{l^{2}} \left( -4dx^{+}dx^{-} + d\vec{y}^{2} \right) + l^{2} \left( (d\psi + P)^{2} + ds_{\mathbb{CP}^{2}}^{2} \right) \\ F_{5} = \dots \\ & \mathbf{Schrödinger} \\ space-time & ds^{2} = \frac{r^{2}}{l^{2}} \left( -2dudv - \frac{r^{2}}{l^{2}}du^{2} + d\vec{y}^{2} \right) + l^{2} \left( (d\psi + P)^{2} + ds_{\mathbb{CP}^{2}}^{2} \right) \\ F_{5} = \dots \\ & B = -\frac{r^{2}}{l}du \wedge (d\psi + P) \end{aligned}$$
Coordinates that eliminate  $\gamma$   $u = 2\gamma lx^{+}$   $v = x^{-}/(\gamma l)$ 

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$$\begin{aligned} \operatorname{AdS}_{5} \times S^{5} & ds^{2} = \frac{r^{2}}{l^{2}} \left( -4dx^{+}dx^{-} + d\bar{y}^{2} \right) + l^{2} \left( (d\psi + P)^{2} + ds_{\mathbb{CP}^{2}}^{2} \right) \\ F_{5} = \dots \\ F_{5} = \dots \\ space-time \\ F_{5} = \dots \\ F_{5} = \dots \\ F_{5} = \dots \\ F_{5} = \dots \\ B = -\frac{r^{2}}{l} du \wedge (d\psi + P) \\ \hline \\ \operatorname{Coordinates that eliminate } \gamma \\ u = 2\gamma lx^{+} \\ v = x^{-}/(\gamma l) \\ \hline \\ \\ \operatorname{GALILEAN SCALING SYMMETRY} \\ \vec{y} \to \lambda \vec{y} \\ u \to \lambda^{2} u \\ v \to v \\ r \to \lambda^{-1} r \\ [\operatorname{Son, Balasubramanian-McGreevy]} \end{aligned}$$

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Obtain *non-relativistic ground state* via *TsT* of  $AdS_5 \times S^5$ 



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Use TsT transformation to get Schrödinger black hole admitting an extremal limit

Use TsT transformation to get Schrödinger black hole admitting an extremal limit

Charged  $Q \neq 0$  AdS-Reissner-Nordström black hole

$$ds^{2} = \frac{r^{2}}{l^{2}} \left(-fdt^{2} + dx^{2} + dy^{2} + dz^{2}\right) + \frac{l^{2}}{r^{2}} \frac{dr^{2}}{f} + l^{2} \left(-d\psi + P + A)^{2} + ds_{\mathbb{CP}^{2}}^{2}\right)$$

$$F_{5} = (1 + \star) \left[ \left(-\frac{4r^{3}}{l^{4}} dt \wedge dr - \frac{2Q}{l^{4}} \omega_{\mathbb{CP}^{2}}\right) \wedge dx \wedge dy \wedge dz \right]$$

$$f(r) = \left(1 - \frac{r^{2}}{r^{2}}\right) \left(1 + \frac{r^{2}}{r^{2}} - \frac{Q^{2}}{r^{2}_{0}r^{4}}\right) \qquad A = A_{t} dt = \frac{Q}{l^{2}} \left(\frac{1}{r^{2}_{0}} - \frac{1}{r^{2}}\right) dt$$

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Charged AdS black hole

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**CONCLUSIONS** 

Use TsT transformation to get Schrödinger black hole admitting an extremal limit

Charged  $Q \neq 0$ AdS-Reissner-Nordström black hole

$$ds^{2} = \frac{r^{2}}{l^{2}} \quad \left(-fdt^{2} + dx^{2} + dy^{2} + dz^{2}\right) + \frac{l^{2}}{r^{2}} \frac{dr^{2}}{f} \\ + l^{2} \left( (d\psi + P + A)^{2} + ds_{\mathbb{CP}^{2}}^{2} \right) \\ F_{5} = (1 + \star) \left[ \left( -\frac{4r^{3}}{l^{4}} dt \wedge dr - \frac{2Q}{l^{4}} \omega_{\mathbb{CP}^{2}} \right) \wedge dx \wedge dy \wedge dz \right] \\ \textbf{DOUBLE ZERO AT HORIZON} \\ f(r) = \left( 1 - \frac{r^{2}_{0}}{r^{2}} \right) \left( 1 + \frac{r^{2}_{0}}{r^{2}} - \frac{Q^{2}}{r^{2}_{0}r^{4}} \right) \quad A = A_{t} dt = \frac{Q}{l^{2}} \left( \frac{1}{r^{2}_{0}} - \frac{1}{r^{2}} \right) dt$$

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Charged AdS black hole

 $f(r) = \left(1 - \frac{r_0}{r^2}\right) \left(1 + \frac{r_0}{r^2} - \frac{\alpha}{r_0^2 r^4}\right)$ 

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Black hole thermodynamics mostly inherited via TsT transformation

 
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Black hole thermodynamics mostly inherited via TsT transformation

Care needed with the identification of CFT time

HOWEVER

[Herzog-Rangamani-Ross]

Black hole thermodynamics mostly inherited via TsT transformation

Care needed with the identification of CFT time

[Herzog-Rangamani-Ross]

Introduce coordinates that *eliminate*  $\gamma$  from *asymptotic Schrödinger metric* 

HOWEVER

$$u = \gamma l (t + x)$$
  $v = \frac{1}{2\gamma l} (t - x)$ 

 
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Partition function  $Z(T, \mu_i) = e^{-W(T, \mu_i)/T} \simeq e^{-I_E}$ 













AGREES WITH  $\gamma$  - INDEPENDENT HORIZON AREA





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# **Extremal black hole and near-horizon limit**

**Extremal limit**  $Q = \sqrt{2} r_0^3$  corresponding to **zero temperature** Admits **near horizon limit** [Reall, Kunduri-Lucietti-Reall]

 
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# **Extremal black hole and near-horizon limit**

**Extremal limit**  $Q = \sqrt{2} r_0^3$  corresponding to **zero temperature** Admits **near horizon limit** [Reall, Kunduri-Lucietti-Reall] **Does an** AdS<sub>2</sub> **arise?** 





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Schrödinger soliton obtained by double Wick rotation of Q = 0 black hole (or TsT of AdS soliton [Horowitz-Myers])

$$t_b \to i x_s \qquad x_b \to i t_s \qquad \gamma_b \to i \gamma_s$$

 
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Schrödinger soliton obtained by double Wick rotation of Q = 0 black hole (or TsT of AdS soliton [Horowitz-Myers])

$$t_b \to ix_s \qquad x_b \to it_s \qquad \gamma_b \to i\gamma_s$$

PERIODICALLY IDENTIFIED FOR REGULARITY







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# **Conclusions and outlook**

In summary	
	We derived a <i>charged black hole</i> solution
	with <i>Schrödinger asymptotic symmetry</i>
	dual to a <i>non-relativistic</i> (2+1)-dimensional CFT at finite temperature /
	finite chemical potential that admits an extremal limit
	The <i>near-horizon geometry</i> exhibits an $\mathrm{AdS}_2$ <i>factor</i> (or <i>not</i> )
	We studied <i>thermodynamic / hydrodynamic</i> properties of the solution
	We found a zero temperature phase transition between black hole and Schrödinger soliton

# **Conclusions and outlook**



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Summary slides follow

### Introduction



### **Motivations**

Study non-relativistic systems at finite temperature / density AdS/CFT Black holes with non-relativistic asymptotic symmetry



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D PHASE LE TRANSITION  $k_F$ 

#### Non-relativistic AdS/CFT



### Non-relativistic AdS/CFT

Obtain *non-relativistic ground state* via *TsT* of  $AdS_5 \times S^5$ 

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### **Charged Schrödinger black hole**



# **Charged Schrödinger black hole**



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### **Thermodynamics**





