

The Kerr/CFT Correspondence

Andy Strominger

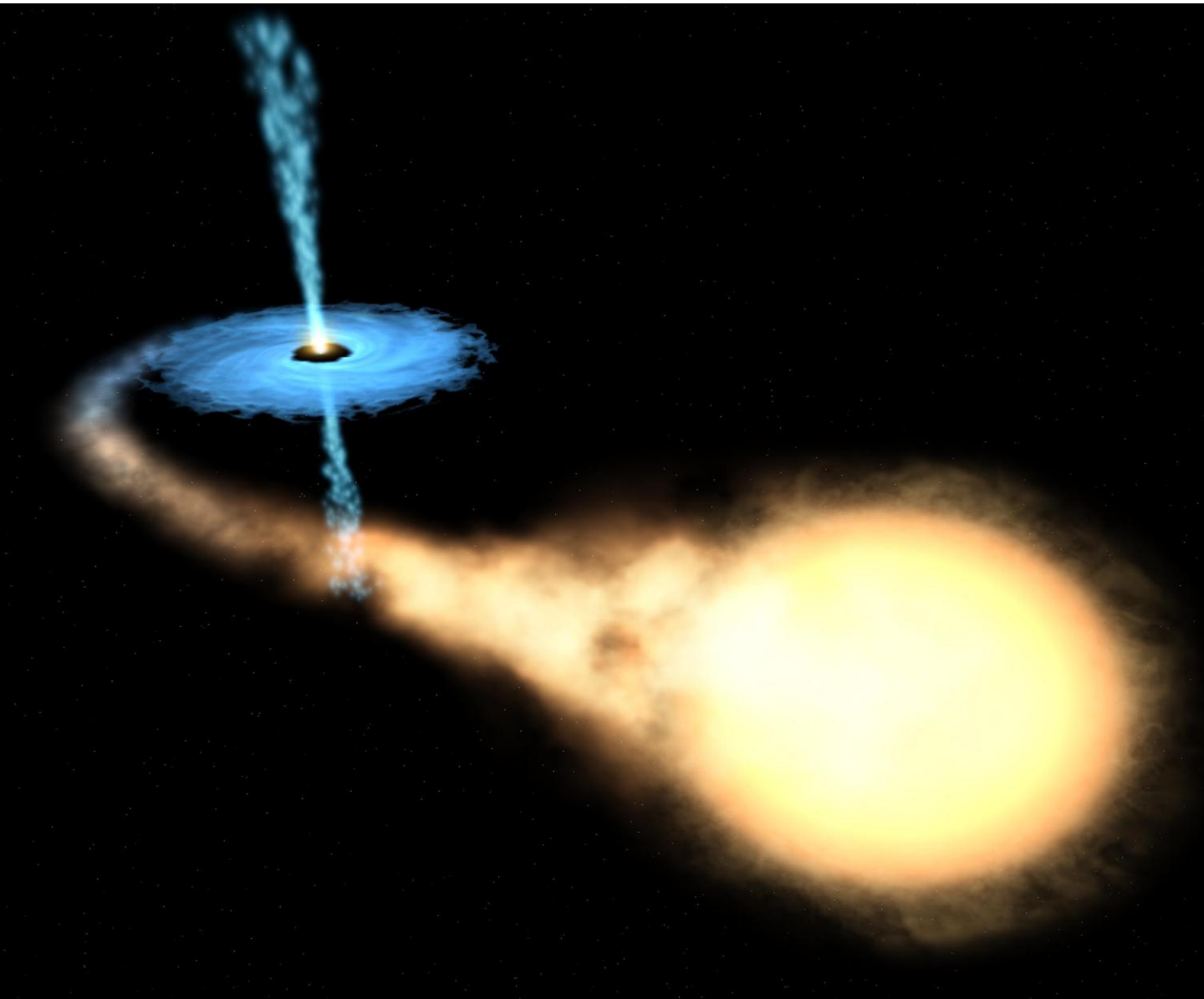
ZURICH 2009

ETH STRING WORKSHOP

w/ M. Guica, T. Hartman
& W. Song I. Bredberg

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In this talk I
will argue & give evidence
that observed extreme Kerr
black holes such as
GRS 1915+105 are dual to
2D CFTs.



An extreme 4D

Kerr  black hole

has angular momentum

$$J = GM^2$$

Hawking temperature

$$T_H = 0,$$

and Bekenstein-Hawking entropy

$$S_{BH} = \frac{2\pi J}{\hbar}.$$

GRS 1915+105 has

$$M \sim 14 M_{\odot}$$

$$\frac{J}{GM^2} > 0.98$$

McClintock, Shafee, Narayan,
Remillard, Davis & Li (2006)

The basic idea is that⁴

Kerr = CFT follows in a near-horizon scaling limit in the spirit of Brown & Henneaux from properties of diffeos.

Condensed Matter

complex
molecular mess

low energy = high redshift = near-horizon
scaling

critical CFT

Astrophysics

complex
binary system

near-horizon
scaling

critical CFT

NASA has kindly prepared a video of our scaling limit. . . .



The near-horizon region
of extreme Kerr



is like chiral gravity:
all excitations must move
counter clockwise at the
speed of light!
So in the extreme case
we get chiral (half of)
2D CFT on (ϕ, t) cylinder.

Role of string theory?

None (except inspirational).
Deriving the universal BH area-entropy law via an exact construction of stringy BH microstates is like deriving the laws of thermodynamics via the periodic table. Boltzmann needed only to assume a consistent UV cutoff (molecules) existed. Similarly, deriving the area law should not require detailed microphysics. We will see it doesn't.

Extreme Kerr Review 7

$$J = M^2 \text{ cosmic censorship}$$

$$ds^2 = - \frac{(\hat{r} - M)^2}{\rho^2} (d\hat{t} - \sin^2 \theta d\hat{\phi})^2 + \frac{\sin^2 \theta}{\rho^2} ((\hat{r}^2 + M^2)d\hat{\phi} - M d\hat{t})^2$$

$$+ \frac{\rho^2}{(\hat{r} - M)^2} d\hat{r}^2 + \rho^2 d\theta^2$$

$$\rho^2 \equiv \hat{r}^2 + M^2 \cos^2 \theta$$

Thermodynamic quantities

$$T_H = 0$$

$$\Omega_H = \frac{1}{2M}$$

$$S_{BH} = \frac{\text{Area}}{4} = 2\pi J$$

EXPLAIN THIS!

Near Horizon Extreme Kerr

Bardeen & Horowitz '99

SCALING

Take $t = \frac{\lambda \hat{t}}{2M}$, $r = \frac{\hat{r} - M}{\lambda M}$, $\phi = \hat{\phi} - \frac{\hat{t}}{2M}$

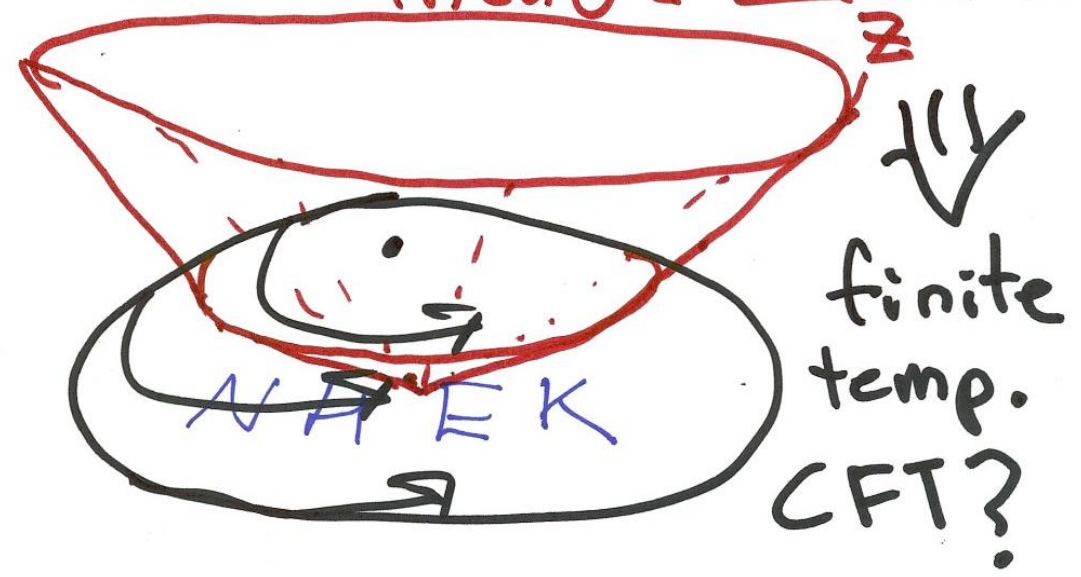
$\lambda \rightarrow 0$ w/ (t, r, θ, ϕ) fixed.

$\Rightarrow ds^2 = 2J\Omega^2 \left[-r^2 dt^2 + \frac{dr^2}{r^2} + d\theta^2 + \Lambda^2 (d\phi + r dt)^2 \right]$

$\Omega^2 \equiv \frac{1 + \cos^2 \theta}{2}$
 $\Lambda \equiv \frac{2 \sin \theta}{1 + \cos^2 \theta}$

ISOMETRY = SVZR/UCI

fixed $\theta =$ Warped AdS₃



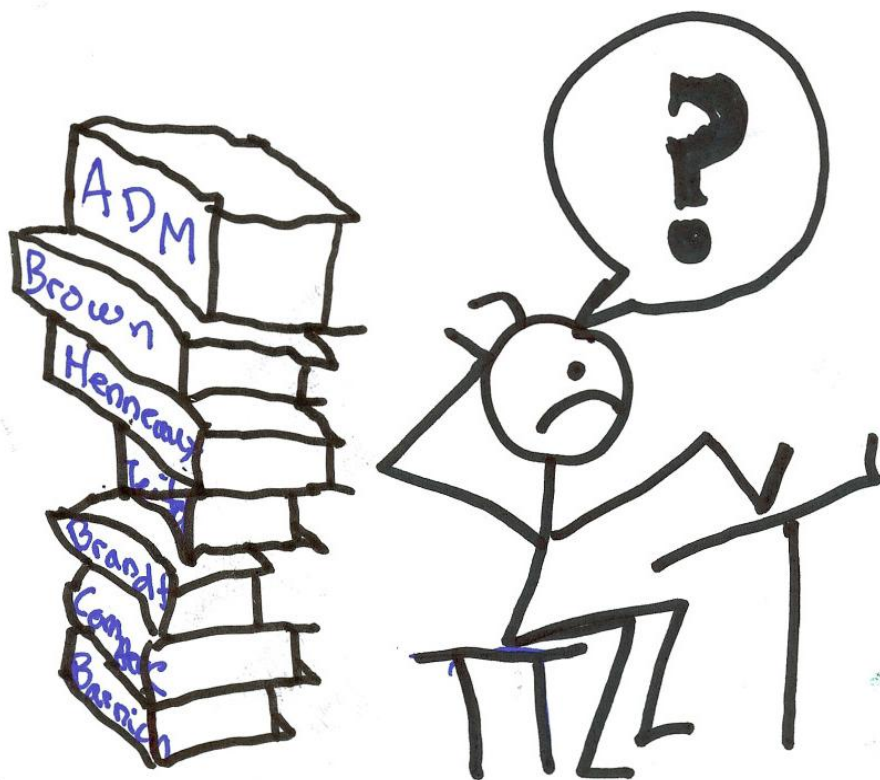
ADS IN THE SKY!!!

H O M E W O R K

Problem 1. Find consistent NHEK boundary conditions.

Problem 2. Find conserved charges.

Problem 3. Determine asymptotic symmetry group.



Asymptotic

Symmetry

Group

depends on boundary conditions

allowed diffeos
trivial diffeos

depends on dynamics

One consistent solution

$$h_{ab} = g_{ab} - g_{ab}^{NHK}$$

$$h_{tt} \sim r^3, h_{t\phi} \sim r^0, h_{t\theta} \sim \frac{1}{r}, h_{tr} \sim \frac{1}{r^2}$$
$$h_{\phi\phi} \sim r^0, h_{\phi r} \sim \frac{1}{r}, h_{\theta\theta} \sim \frac{1}{r}, h_{\theta r} \sim \frac{1}{r^2}$$
$$h_{rr} \sim \frac{1}{r^3}$$

PLUS $M^2 = 5$ Divide & STUDY conquer) EXTRA
charges are finite of allowed ONLY diffeos, not all zero.

SIMILAR TO BMS GROUP.

CENTRAL CHARGE

The ASG for NHEK is generated by

$$S_n = -e^{-in\phi} \partial_\phi + ine^{-in\phi} r \partial_r$$

w/ Lie Bracket algebra = Virasoro

$$i [S_m, S_n]_{L.B.} = (m-n) S_{m+n}$$

The action on fields is generated via Dirac Brackets w/ the conserved charges

$$Q_n[h] = -\frac{1}{32\pi G} \int_{S^2} * dx^\mu \wedge dx^\nu [S_{n\nu} D_\mu h$$

Abbot Deser, Barnich Brandt, Compere, Brown Henneaux...

$$- S_{n\nu} D^\nu h_{\mu\sigma} + S_n D_\nu h_{\mu\sigma} + \frac{\hbar}{2} D_\nu S_{\mu\sigma}$$

$$- h_{\nu\sigma} D^\sigma S_n + \frac{1}{2} \hbar \delta_{\nu\sigma} (D_\mu S_n^\sigma + D^\sigma S_{n\mu})$$

THESE ARE FINITE

By construction

$$i \{Q_m, Q_n\}_{D.B.} = (m-n) Q_{m+n} + i Q_n [L, S_m]_{NHEK}$$



$$C_L = 12 J$$

This implies that if we succeed in constructing QG of any kind on NHEK with the given BC it must be a 2D CFT in the sense that the quantum states form Virasoro reps. Bulk locality \leadsto boundary CFT locality, maybe. Physical observation suggests the CFT is unitarity.

So far, we know only the central charge of the CFT. Physical observation \rightarrow properties of CFT. Knowing the CFT exactly = knowing all the laws of physics. We don't expect that just yet!

THE ANALYSIS IS BOTTOM UP, NOT TOP DOWN!

TEMPERATURE

A general Kerr black hole corresponds to the thermodynamic ensemble:

$$\rho = e^{-\frac{(\omega - m\Omega_H)}{T_H}}$$

where ω, m are Fourier coefficients.

Define n_R, n_L by

$$e^{-i\omega t + im\phi} = e^{-in_R t + in_L \phi}$$

\Rightarrow $n_L = m = L_0$ = angular momentum
 $n_R = \frac{1}{\lambda}(2\omega - m)$ = energy above extremality

Define temps T_L, T_R conjugate to n_L, n_R

$$\rho = e^{-\frac{n_R}{T_R} - \frac{n_L}{T_L}}$$

For extremality $T_H \rightarrow 0, n_R$

$T_L \rightarrow \frac{1}{2\pi}$, $T_R \rightarrow 0$

Ensemble of left movers

$$\rho = e^{-2\pi n_L}$$

Entropy

The canonical Cardy formula is

$$S = \frac{\pi^2}{3} c_L T_L$$

Using

$$c_L = 12J, \quad T_L = \frac{1}{2\pi}$$

gives

$$S_{\text{micro}} = 2\pi J = \frac{\text{Area}}{4} !$$

Summary

Canonical symmetry analysis \Rightarrow extreme Kerr is dual to a $c = 12J$ CFT at $\mathbb{F} = \frac{1}{2\pi}$. Assuming Cardy, unitarity etc, this explains extreme Kerr entropy.

What's new/next?

The analysis can be generalized to higher D , charges, c.c., multiple angular momenta etc. It always works. Agreement much more non-trivial

horan Soltanpanahi Lu Mei Pope Ogawa
 Azeyanagi Terashima Murata Nishioka
 Hartman AS Peng Wu Isono Tai Wen
 Chow Nakayama Chen Wang Compere
 Vasquez-Poditz Wu Tian Tachikawa
 Carolus Ghodsi Matsuo Tsukioka Yoo

Lesson \mathbb{Z} inequivalent consistent
 BCs w/ different dual CFTs, T_h, C_L
 but same SBH.

Higher derivative corrections to Wald
 also matched. Krishnan Kuperstein
 Azeyanagi Compere Ogawa Terashima
 Tachikawa

Hints: Kerr Horizon = Surface of Fermi Sea?

Emparan et al Verlinde et al
 McGreevy Lu et al

Classical GR / dynamics
of NHEK largely unexplored
beyond Bardeen & Horowitz.

- What are possible b.c.s?
- What is causal structure/geodesics?
- Positive energy theorem, uniqueness?
- Instabilities?
- How do we describe the boundary?
- ⋮
- !

Recent progress

Amsel, Horowitz, Marolf, Roberts I, II
 Dias, Reall Santos
 Balasubramanian, deBoer, Sheikh-Jabbari,
 Simon

Seems to confirm uniqueness of NHEK
up to diffeos (which may be nontrivial!)

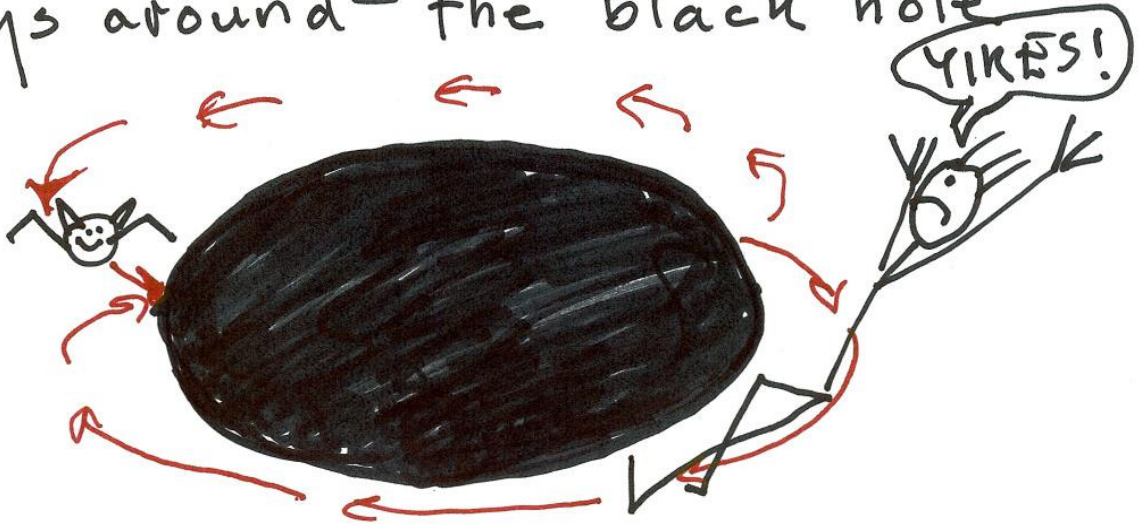
EVENTUALLY KERR/CFT MAY EXPLAIN CURRENTLY MYSTERIOUS OBSERVED BLACK HOLE BEHAVIORS.

BEYOND EXTREMALITY

"BLACK HOLE SUPERRADIANCE FROM KERR/CFT"

J. Bredberg, T. Hartman, W. Song
& A.S.†

For $M^2 > J$ things can move both ways around the black hole



so we expect a non-chiral CFT, with \bar{L}_n enhancing $SL(2, \mathbb{R})$ of NHEK.

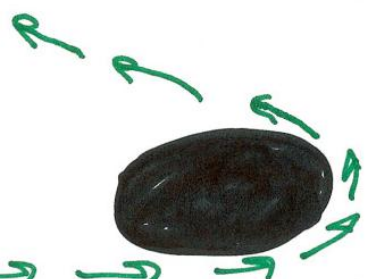
Matsuo Tsukioka Yoo

C_R=12J!

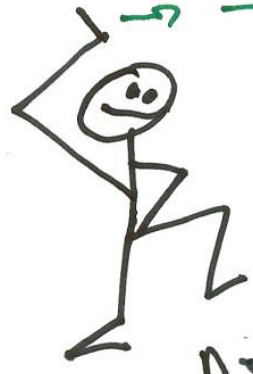
Castro Larsen

This is a hard problem, due to AdS_2 -type IR divergences. We have taken the smallest step away from extremality we could find....

Scattering Super radiant



Do ~~CFT~~ and gravity agree?
c.f. Maldacena #AS '96



Kerr scattering cross sections for massless scalar were computed in Starobinsky & Churilov '73, Teukolsky & Press '74. In the near-'superradiant' limit

$\Gamma_R \rightarrow 0$, $\tilde{n}_R = \frac{n_R}{2\pi\Gamma_R}$ fixed, their formula is

$$G_{abs} = \frac{|\Gamma(\frac{1}{2} + \beta + i\tilde{n}_R)|^2 (m^2 \alpha M^2 \hbar)}{2\pi \Gamma(2\beta)^2 \Gamma(2\beta + 1)^2} \times e^{-\pi \tilde{n}_R} |\Gamma(\frac{1}{2} + \beta + i\tilde{n}_R)|^2 \times \Gamma|\frac{1}{2} + \beta + i\tilde{n}_R|^2$$

2β Normalization from near far matching
 $\langle \mathcal{O}_\beta(0) \mathcal{O}_\beta(x^+) \rangle_{CFT}$
 $\langle \mathcal{O}_\beta(0) \mathcal{O}_\beta(x) \rangle_{CFT}$

$\beta^2 = k_g^2 - 2m^2 + \frac{1}{4}$
 \leftarrow separation constants

Matches perfectly CFT calculation with $\dim(\mathcal{O}) = \frac{1}{2} + \beta$, computed from highest weight solns of wave eq.

ALSO WORKS FOR KERR-NEWMAN, HIGHER SPIN!

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In fact the match between bulk gravity and boundary CFT is better than expected. We don't understand why it works so well! But we seem to be on the right track.

Summary

- 1) Extreme Kerr entropy is matched by $c=125$ chiral CFT.
- 2) Scattering cross sections reproduced by non-chiral CFT.
- 3) Many aspects of Kerr/CFT remain mysterious.